

Measurements and Reaction Modeling for Proton Bombardment on Natural Antimony

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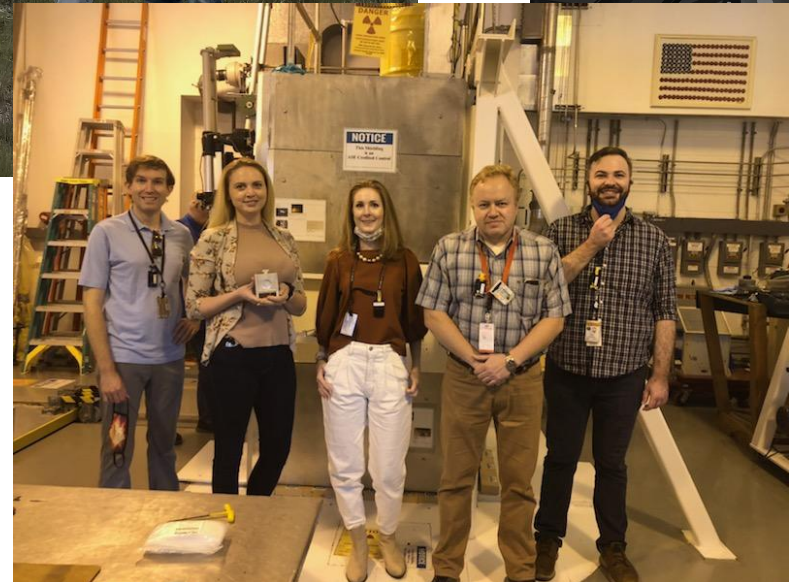
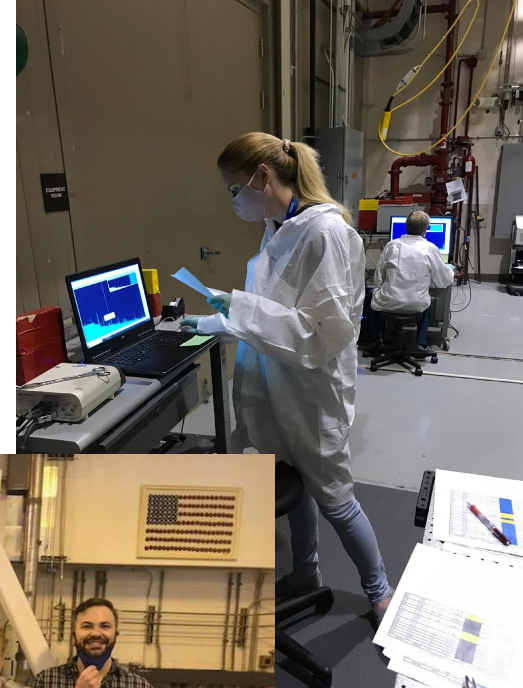
2023 NSSC Fall Workshop
Berkeley, CA

University Affiliation:
University of California, Berkeley

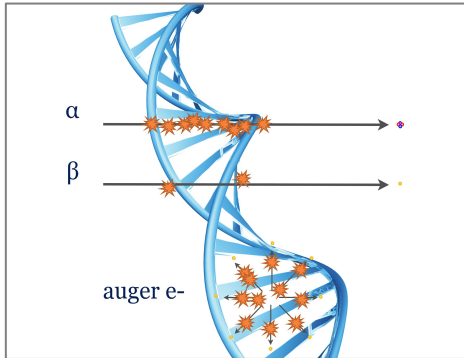
Standing:
PhD Candidate, Nuclear Engineering

Research Focus:
Medical radioisotope production
Nuclear data
Charged particle reaction modeling

Lab Affiliation:
Los Alamos National Lab
Lawrence Berkeley National Lab
Brookhaven National Lab



- Cross section measurements of ^{117m}Sn and ^{119m}Te (a cow for ^{119}Sb) to optimize production for high SA with minimal impurities
 - These products can be produced via $^{nat}\text{Sb}(p,x)$



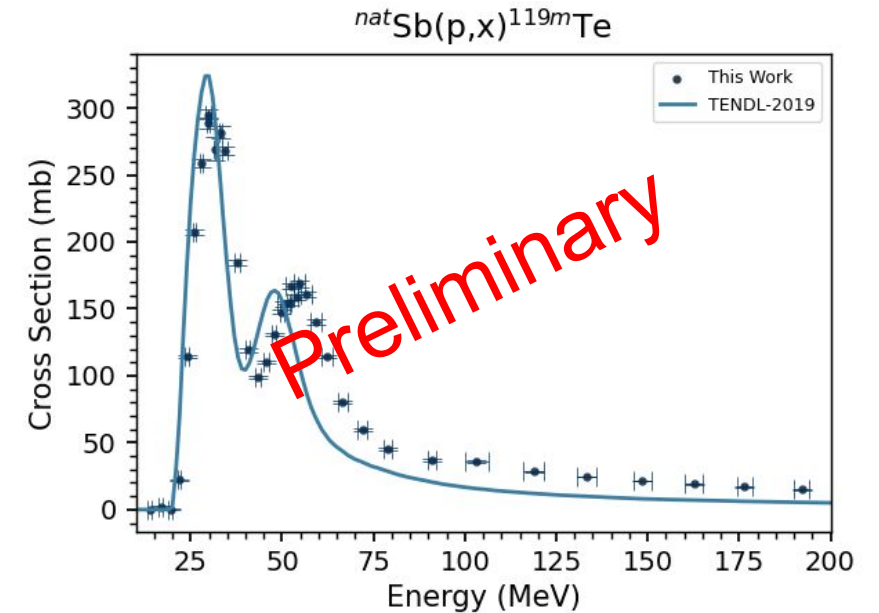
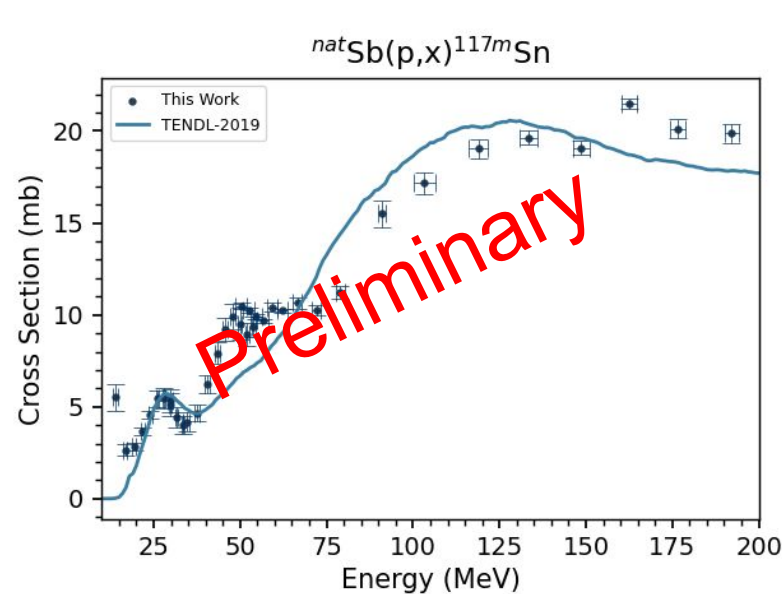
Primary Isotopes:

$^{119m}, ^{119g}\text{Te}, ^{117m}\text{Sn}$

Potential Impurities:

$^{116}, ^{117}, ^{118}, ^{121m}, ^{121g}, ^{123m}\text{Te},$

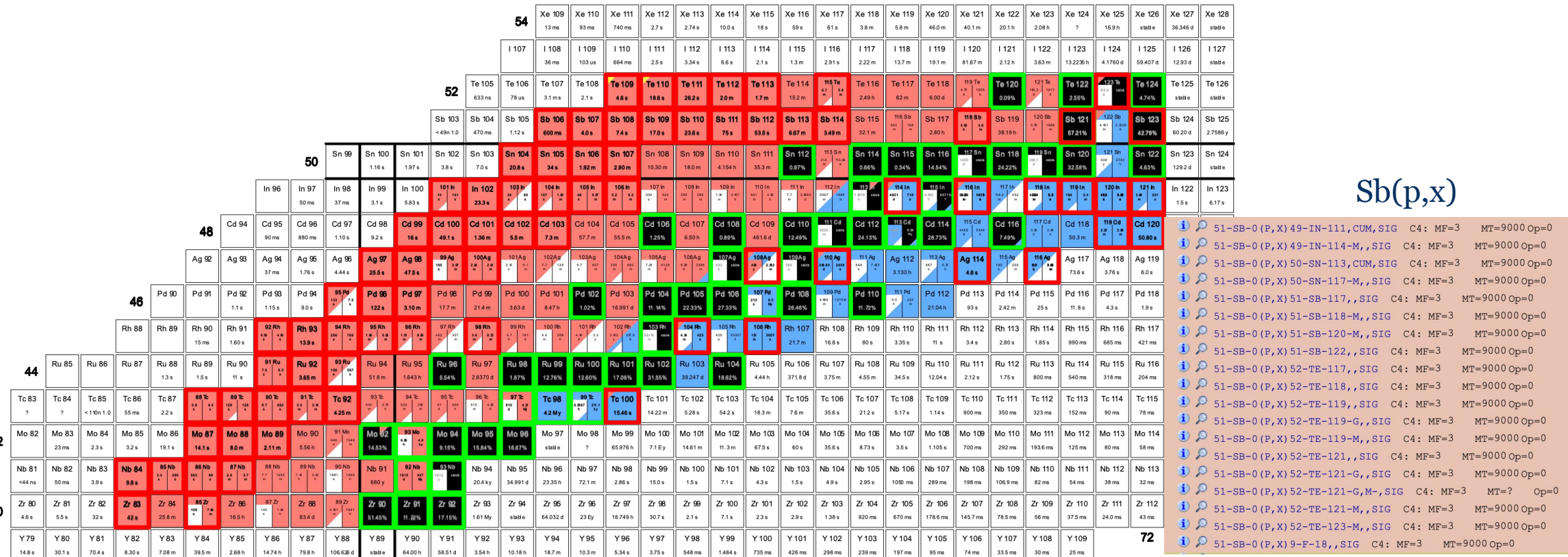
$^{113}, ^{119m}, ^{121m}, ^{121g}\text{Sn}$



This is a compelling motivation, but there's more!

Motivation: Sb(p,x) Reaction Modeling

Incident proton energy up to 200 MeV opens channels for 200+ potential products!



We can use this data to improve our reaction modeling capabilities!

Experimental Setup: Overview

A Tri-lab collaboration has been formed between LBNL, LANL, and BNL to measure (p,x) reactions relevant to isotope production from threshold to 200 MeV *for primary isotopes of interest and their impurities.*



LBNL 88-Inch Cyclotron

$$E_{p,max} = 60 \text{ MeV}$$



LANL IPF

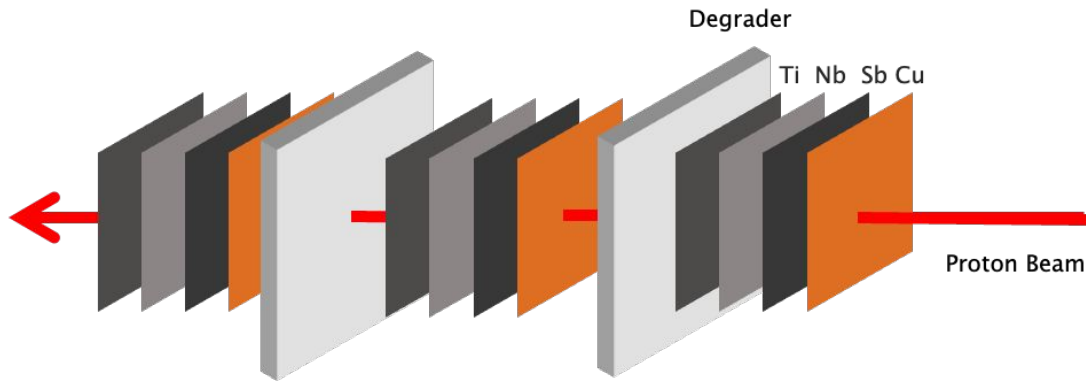
$$E_{p,max} = 100 \text{ MeV}$$



BNL BLIP

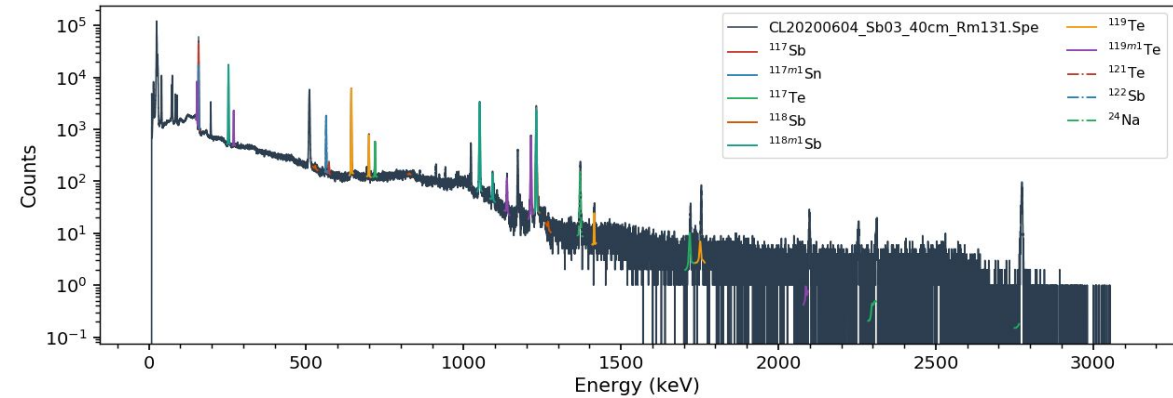
$$E_{p,max} = 200 \text{ MeV}$$

Target stack irradiated with 150 nA proton beam



Beam boxes at LBNL, BNL

Foils removed and counted on HPGe detectors

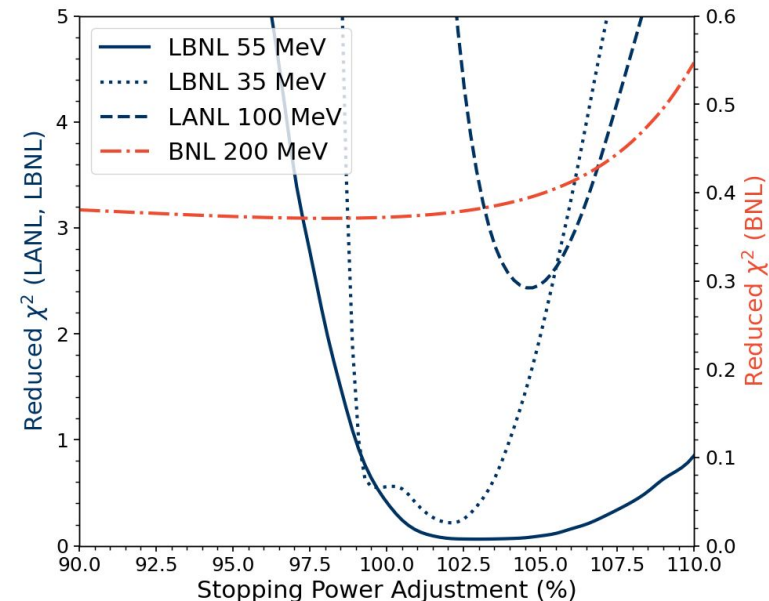
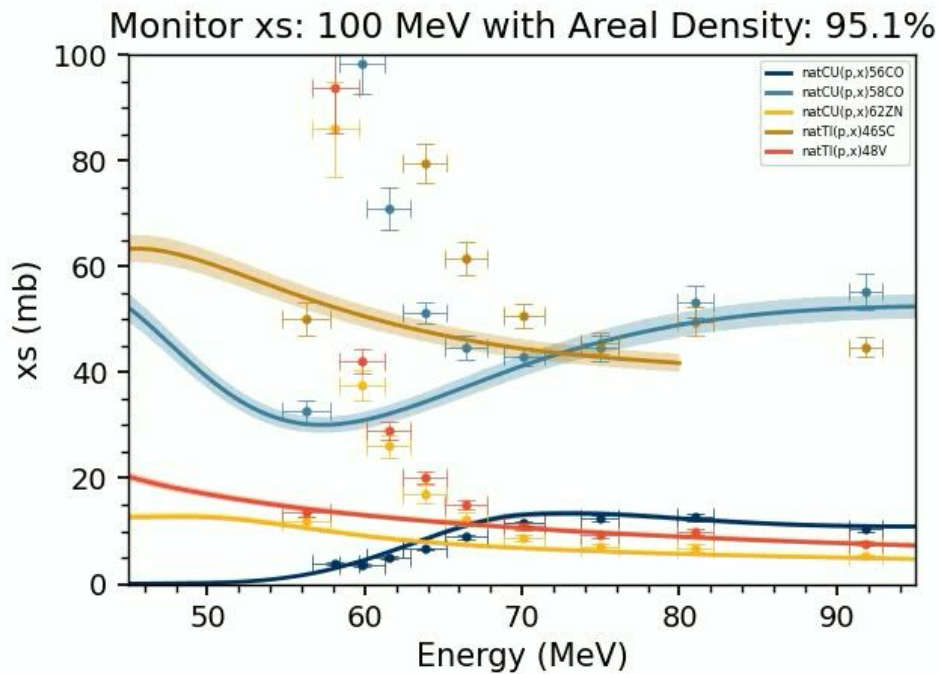


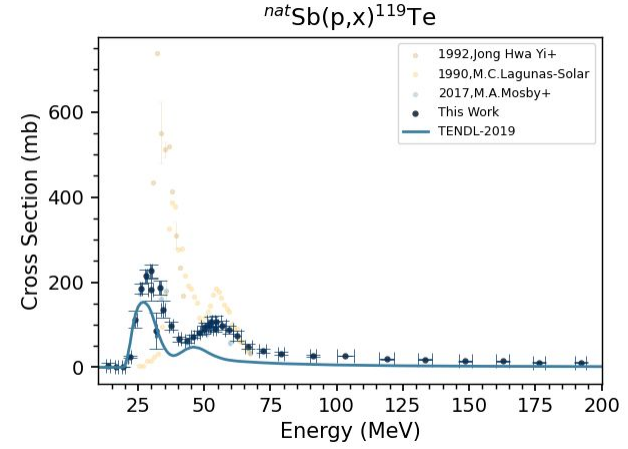
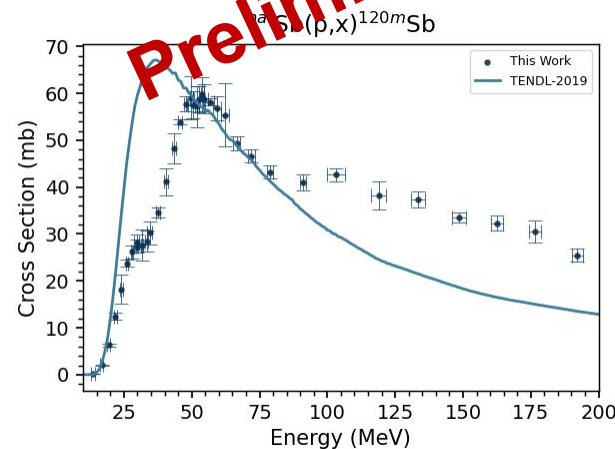
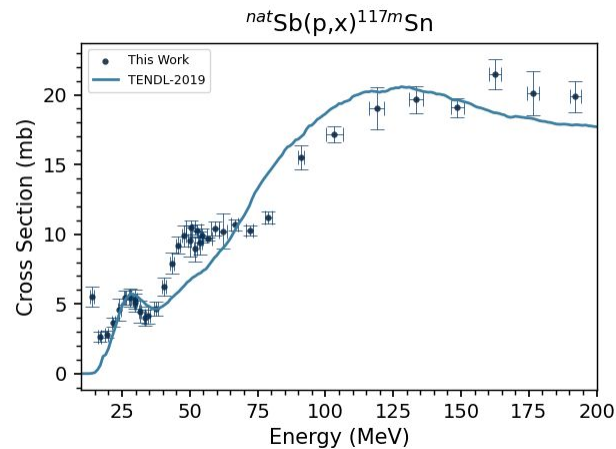
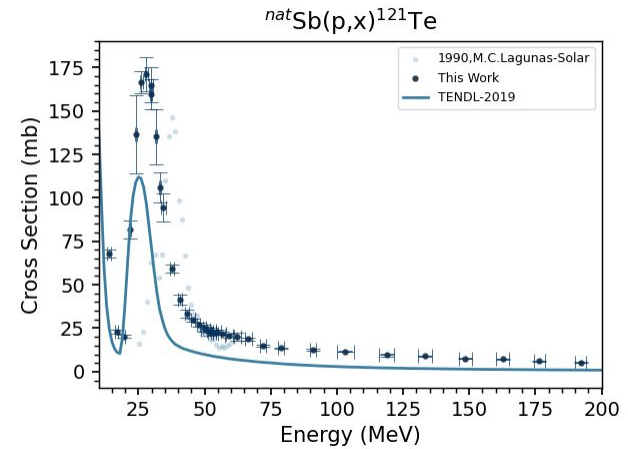
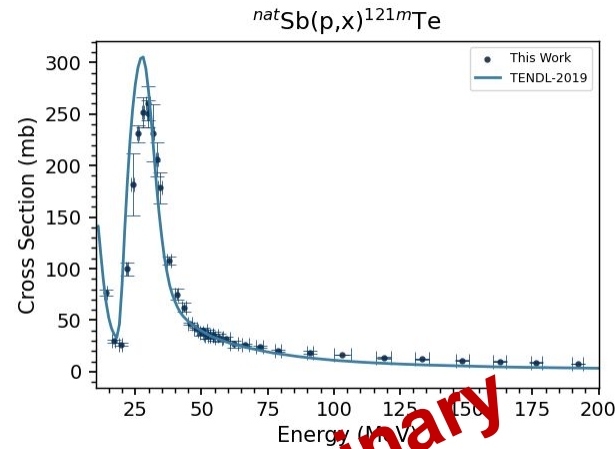
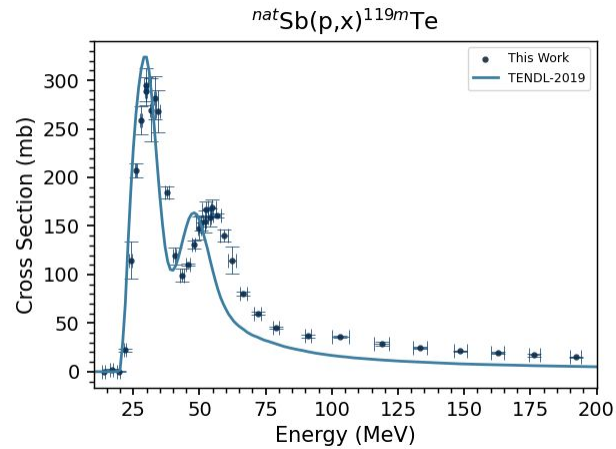
Sample gamma spectrum of irradiated Sb, with products identified

Nominally varying the areal density of the stack materials provides a better representation of current and energy in each bin reduces systematic uncertainty due to range straggling and limitations in stopping power characterization.

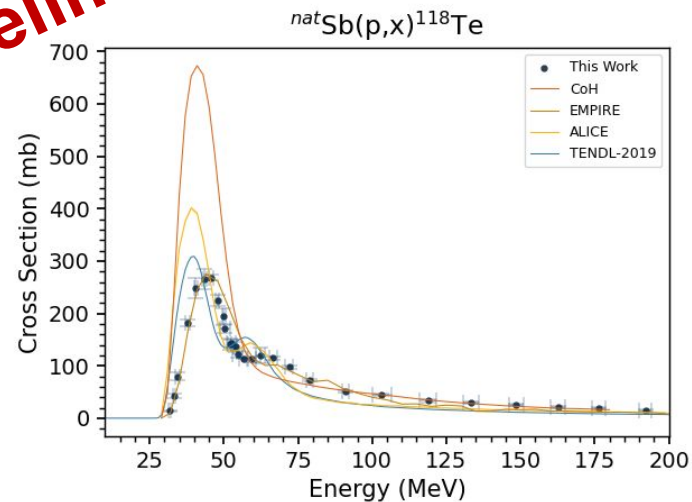
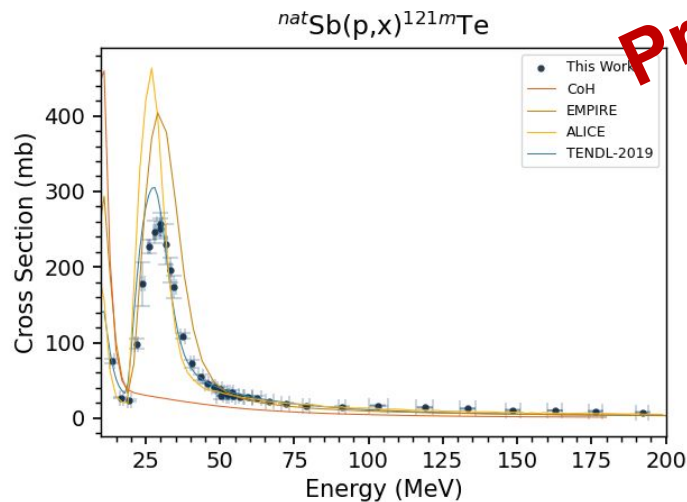
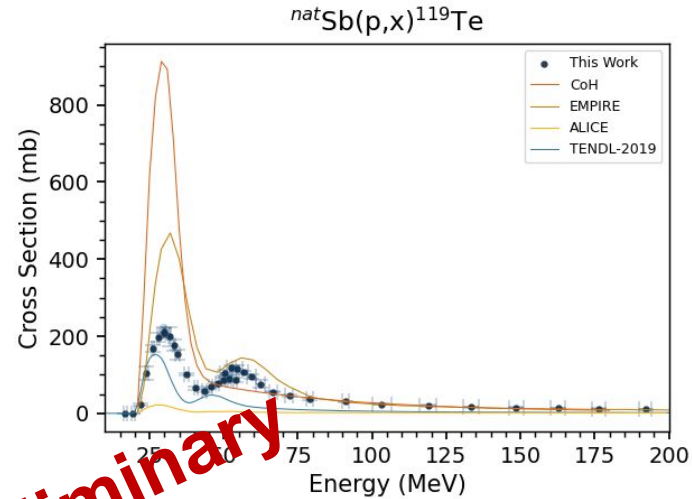
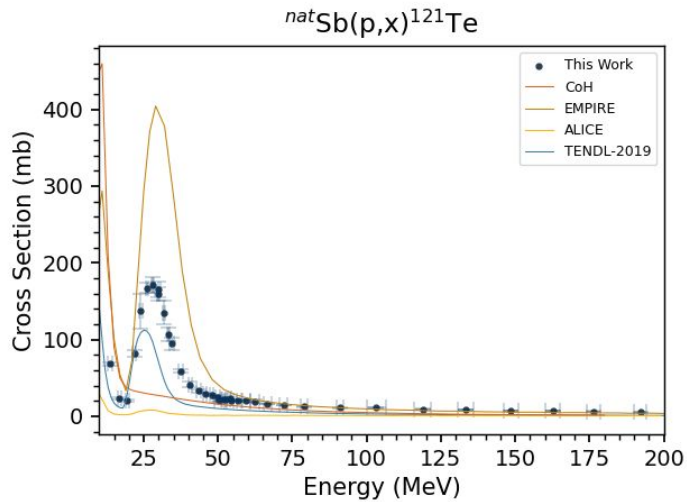
This variance minimization technique has been utilized in *Graves et al.*, *Voyles et al.*, *Fox et al.*, and *Morrell et al.*

$$I = \frac{A_0}{\rho \Delta x (1 - e^{-\lambda t}) \int \sigma(E) \frac{d\phi}{dE} dE}$$





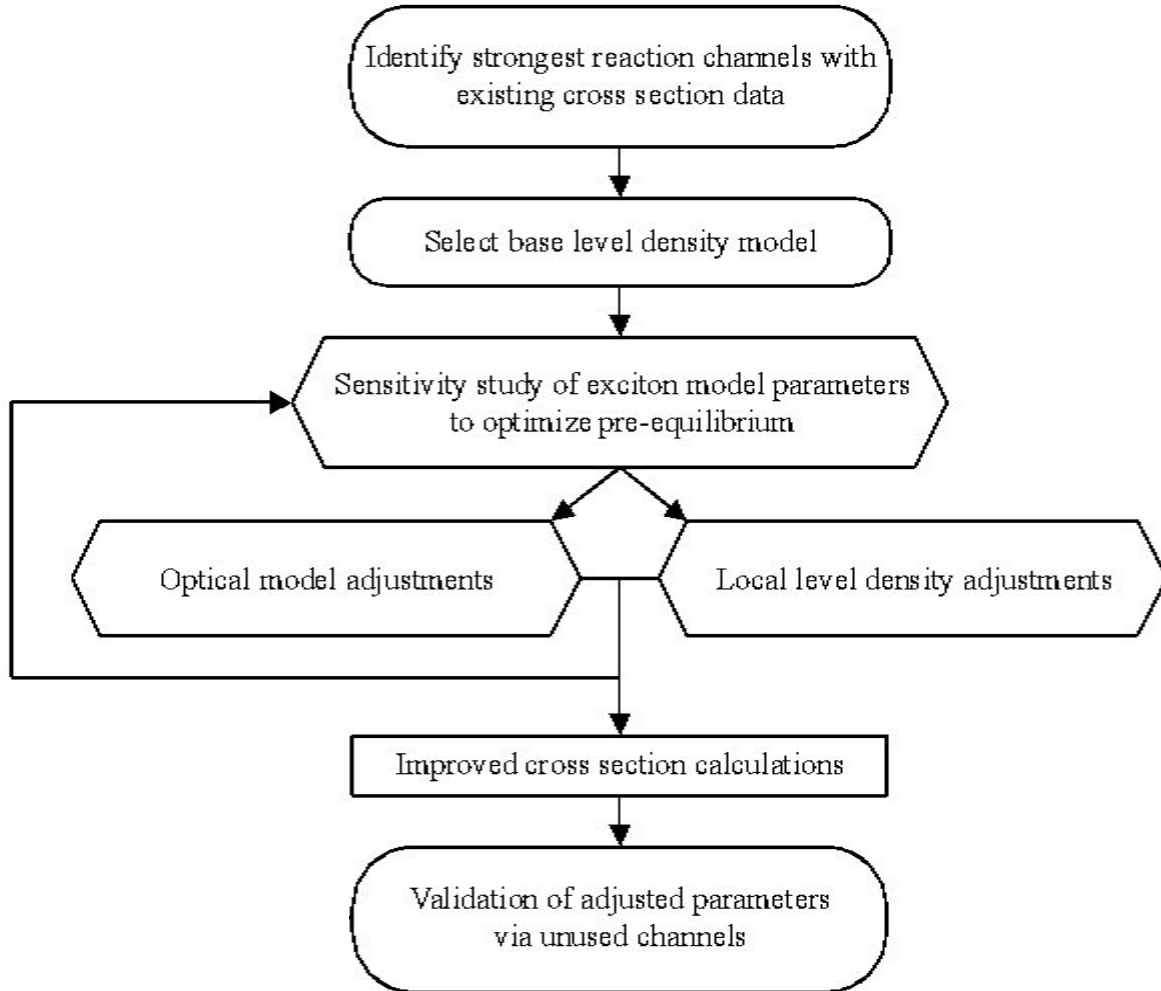
Preliminary



Preliminary

Compared experimental results to standard inputs for:

- ALICE 3.5.3
- CoH 3.2.3
- EMPIRE 3.2.3
- TENDL 2019

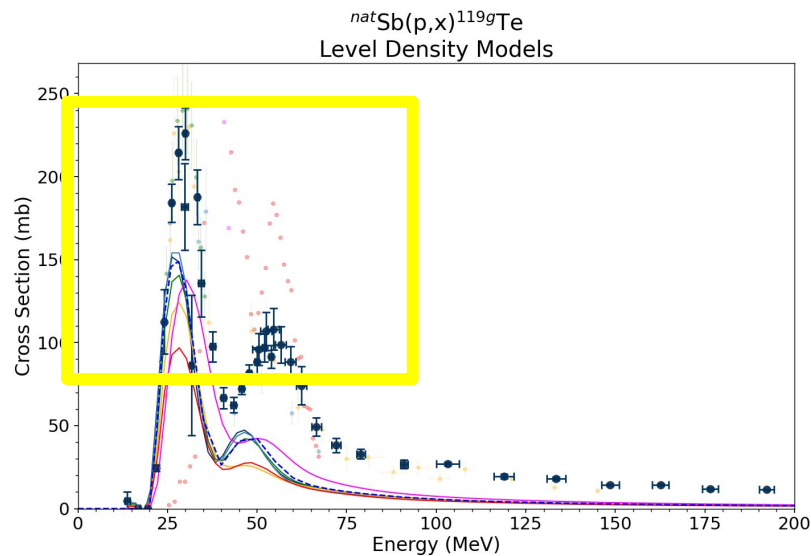
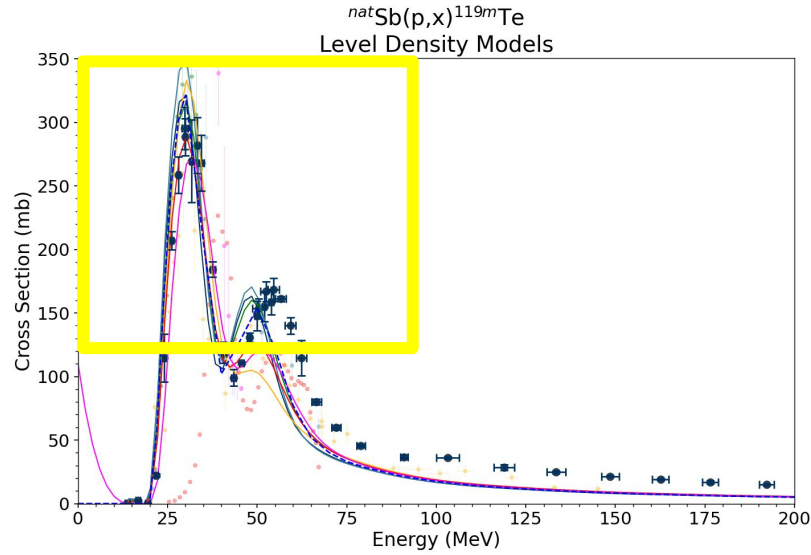


- Iterative approach

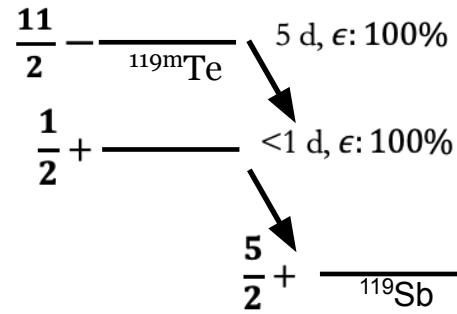
- Select level density model
- Adjust level density models
- Adjust pre-equilibrium parameters
- Adjust optical model potential parameters
- Sensitivity review on previous adjustments

Courtesy Fox et al. 2021

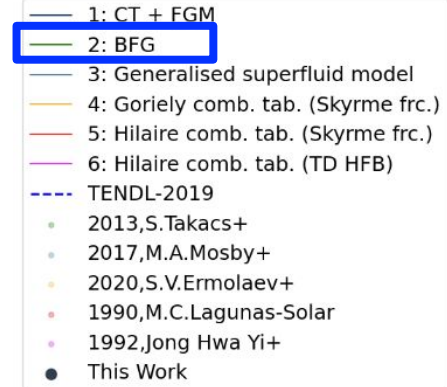
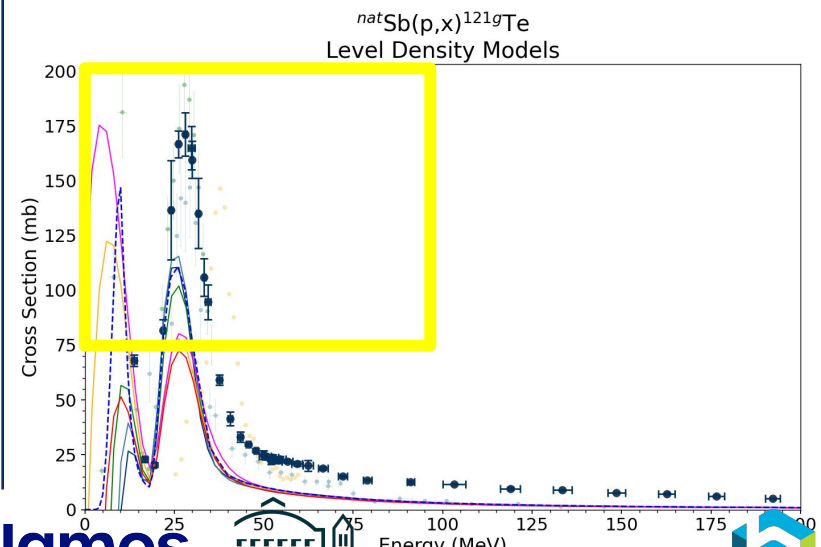
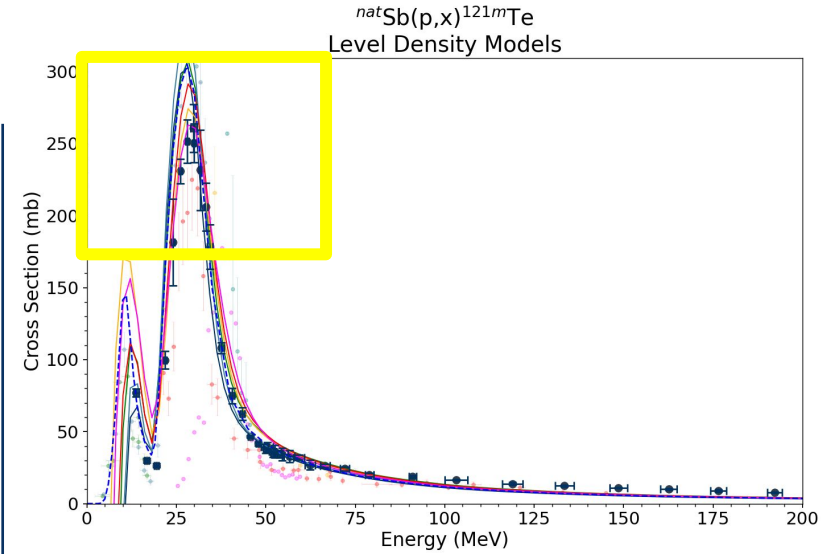
Reaction Modeling: High Spin Isomers



Preliminary

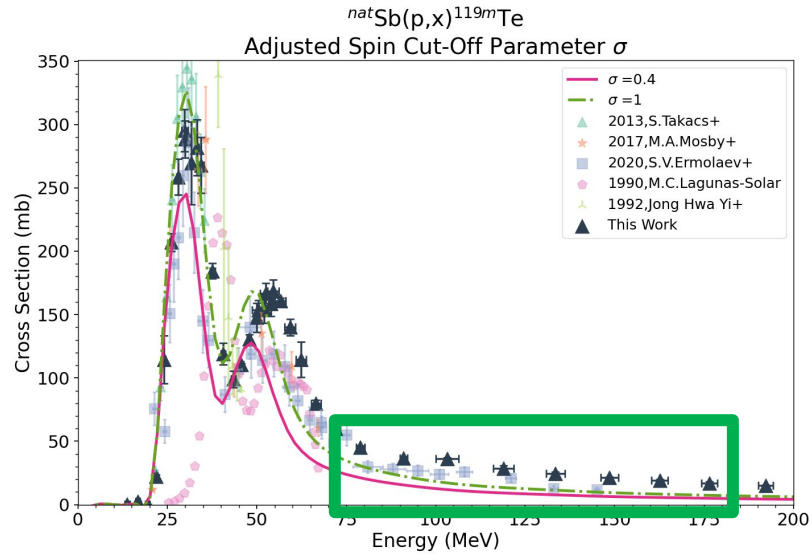


Underprediction
for ground states

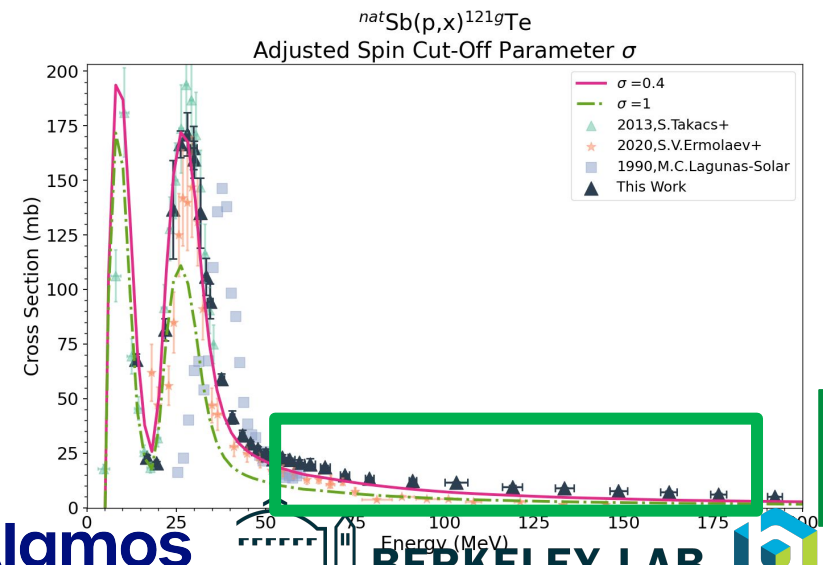
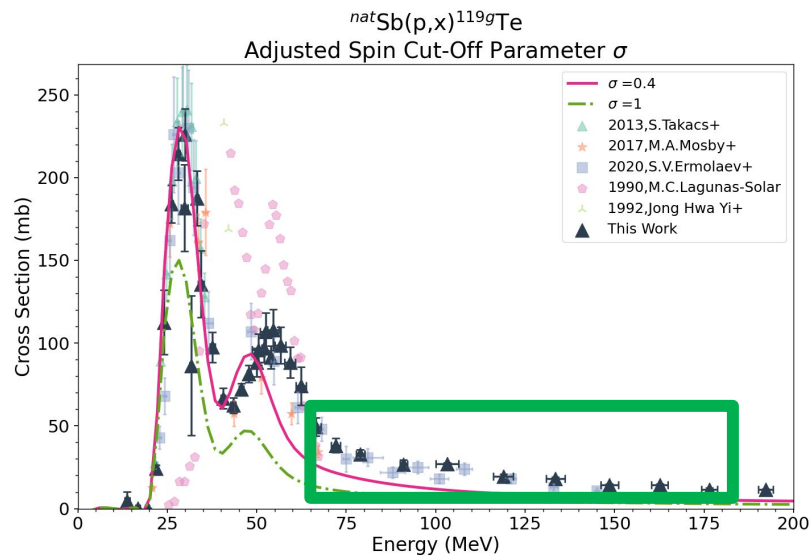
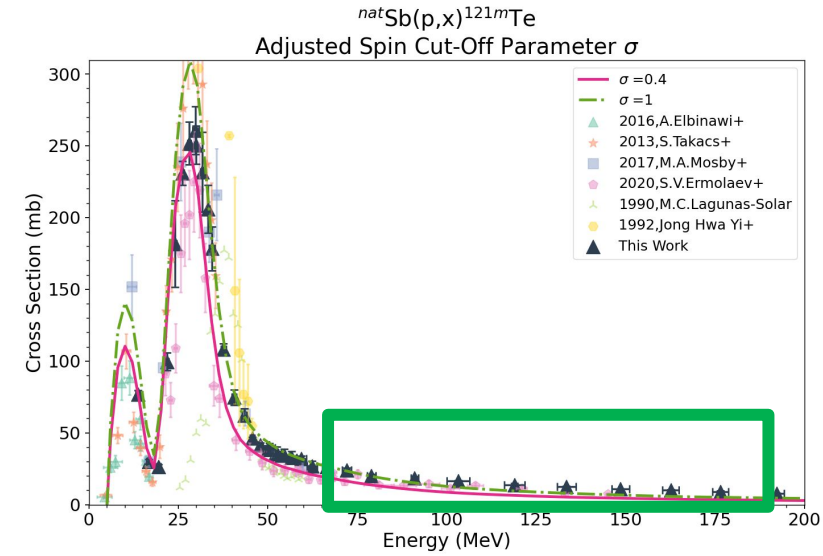


*BFG suggested in
Khandaker et al. 2021*

Reaction Modeling: Adjusting the Spin Cut-off Parameter



Preliminary



Consistent
under-prediction in

Brookhaven 11
National Laboratory

Preequilibrium reaction adjustments -2 component exciton model:

- **M2Constant, M2Limit, M2Shift** previously explored by *Fox et al.* to adjust the effective squared matrix M^2

$$M^2 = \frac{C_1 A_p}{A^3} \left[7.48 C_2 - \frac{4.62 \times 10^5}{\left(\frac{E^{\text{tot}}}{n \cdot A_p} + 10.7 C_3 \right)^3} \right]$$

- **Rpipi, rpinu, rnupi, rnunu** adjustments based on nucleon-nucleon interactions

$$\begin{aligned} M_{\pi\pi}^2 &= R_{\pi\pi} M^2 \\ M_{\nu\nu}^2 &= R_{\nu\nu} M^2 \\ M_{\pi\nu}^2 &= R_{\pi\nu} M^2 \\ M_{\nu\pi}^2 &= R_{\nu\pi} M^2 \end{aligned}$$

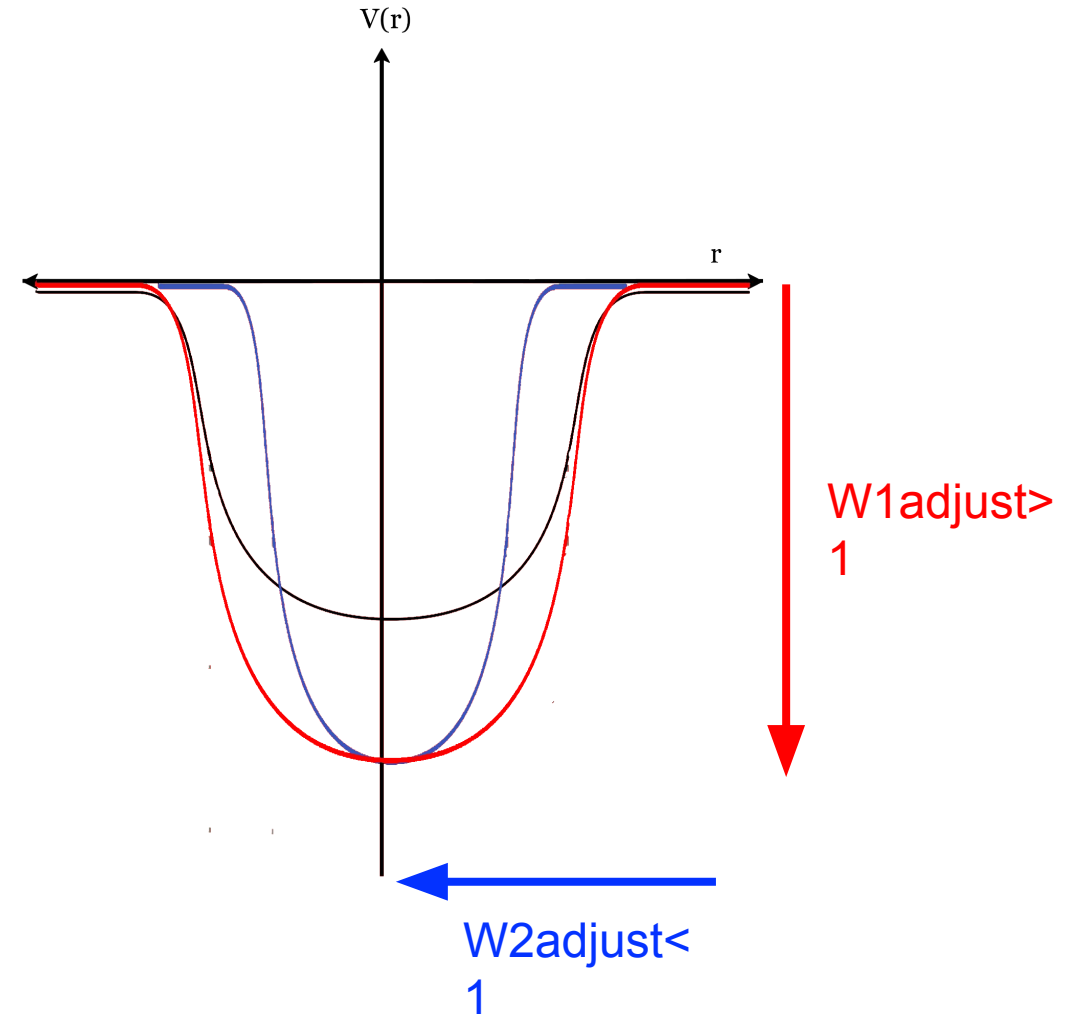
Variable	Default	Range	Adjusted
M2Constant	1	0-100	2
M2Limit	1	0-100	0.8
M2Shift	1	0-100	1.8
	1	0-100	1.5
	1	0-100	1
	1	0-100	1.5
	1.5	0-100	1.5

OMP adjustments to the imaginary volume term:

- **W1adjust**, **w2adjust** adjustments to imaginary volume term

$$W_V(E) = w_1^n \frac{(E - E_f^n)^2}{(E - E_f^n)^2 + (w_2^n)^2}$$

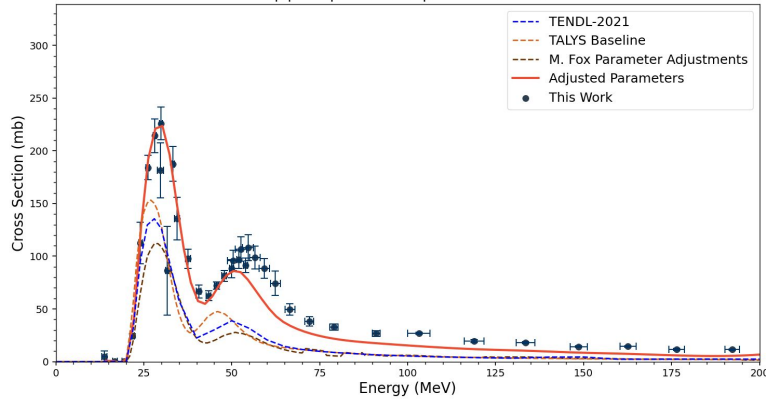
Variable	Default	TALYS Range	Adjusted
W1adjust n	1	0.1-10	2.5
W2adjust n	1	0.1-10	0.6



Reaction Modeling: Results

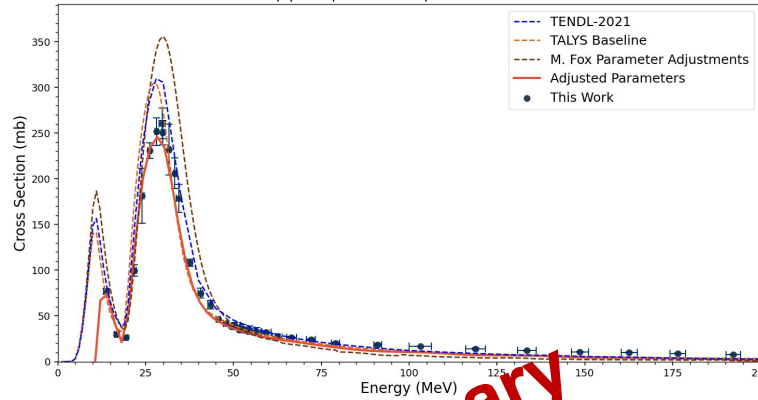
$^{nat}\text{Sb}(p,x)^{119}\text{Te}$

LD Model: 2, Spin Cut-Off Model: 2, Equidistant: y
M2Constant: 2, M2Limit: 0.8, M2Shift1.8
Spin Cut-Off Multiplier: 0.4 Colldamp: n preeqmode: 1 mpreeqmode:2 preeqspin: 1
w1Adjust n: 2.5, w2Adjust n: 0.6
rpi: 1 rpinu: 1.5 rmu: 1.5 rnu: 1.5



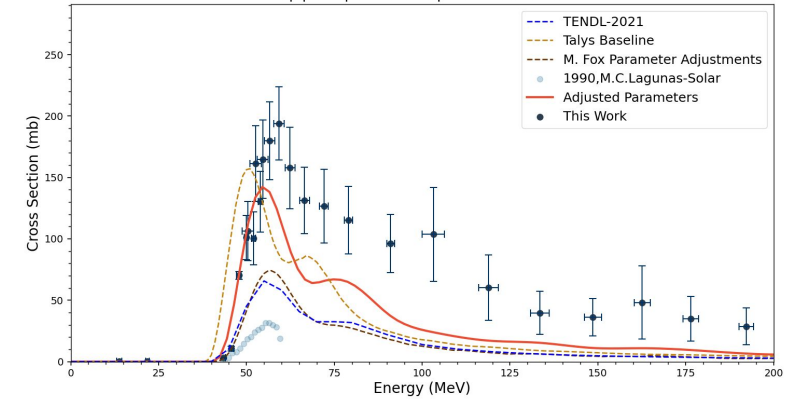
$^{nat}\text{Sb}(p,x)^{121m}\text{Te}$

LD Model: 2, Spin Cut-Off Model: 2, Equidistant: y
M2Constant: 2, M2Limit: 0.8, M2Shift1.8
Spin Cut-Off Multiplier: 0.4 Colldamp: n preeqmode: 1 mpreeqmode:2 preeqspin: 1
w1Adjust n: 2.5, w2Adjust n: 0.6
rpi: 1 rpinu: 1.5 rmu: 1.5 rnu: 1.5



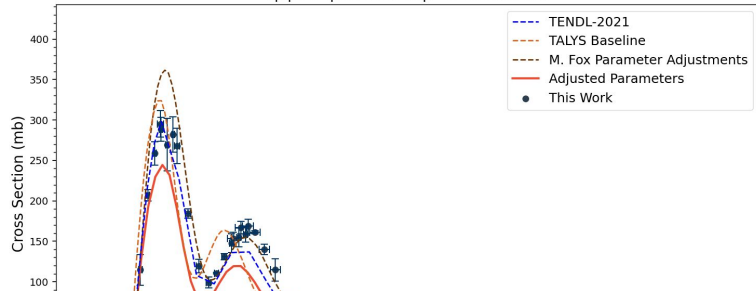
$^{nat}\text{Sb}(p,x)^{117}\text{Te}$

LD Model: 2, Spin Cut-Off Model: 2, Equidistant: y
M2Constant: 2, M2Limit: 0.8, M2Shift1.8
Spin Cut-Off Multiplier: 0.4 Colldamp: n preeqmode: 1 mpreeqmode:2 preeqspin: 1
w1Adjust n: 2.5, w2Adjust n: 0.6, cknock a: 1
rpi: 1 rpinu: 1.5 rmu: 1.5 rnu: 1.5



$^{nat}\text{Sb}(p,x)^{119m}\text{Te}$

LD Model: 2, Spin Cut-Off Model: 2, Equidistant: y
M2Constant: 2, M2Limit: 0.8, M2Shift1.8
Spin Cut-Off Multiplier: 0.4 Colldamp: n preeqmode: 1 mpreeqmode:2 preeqspin: 1
w1Adjust n: 2.5, w2Adjust n: 0.6
rpi: 1 rpinu: 1.5 rmu: 1.5 rnu: 1.5



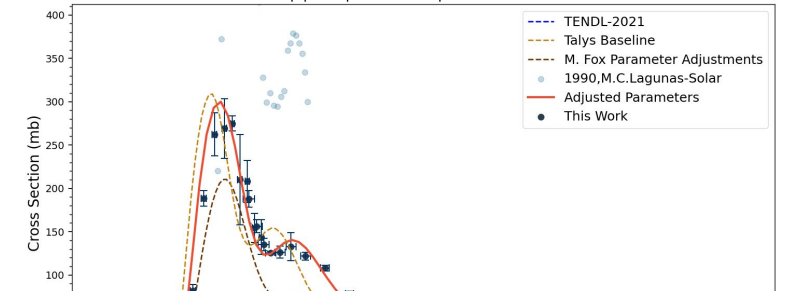
$^{nat}\text{Sb}(p,x)^{121}\text{Te}$

LD Model: 2, Spin Cut-Off Model: 2, Equidistant: y
M2Constant: 2, M2Limit: 0.8, M2Shift1.8
Spin Cut-Off Multiplier: 0.4 Colldamp: n preeqmode: 1 mpreeqmode:2 preeqspin: 1
w1Adjust n: 2.5, w2Adjust n: 0.6
rpi: 1 rpinu: 1.5 rmu: 1.5 rnu: 1.5



$^{nat}\text{Sb}(p,x)^{118}\text{Te}$

LD Model: 2, Spin Cut-Off Model: 2, Equidistant: y
M2Constant: 2, M2Limit: 0.8, M2Shift1.8
Spin Cut-Off Multiplier: 0.4 Colldamp: n preeqmode: 1 mpreeqmode:2 preeqspin: 1
w1Adjust n: 2.5, w2Adjust n: 0.6, cknock a: 1
rpi: 1 rpinu: 1.5 rmu: 1.5 rnu: 1.5

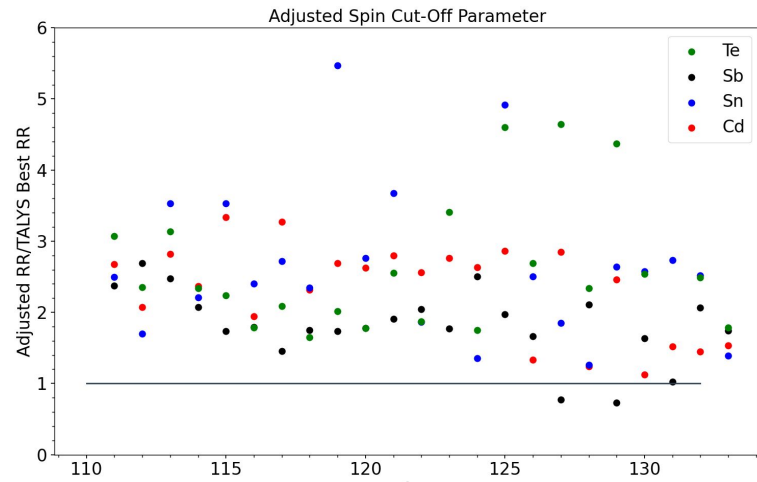


Preliminary

What are the effects of implementing these parameter adjustments?

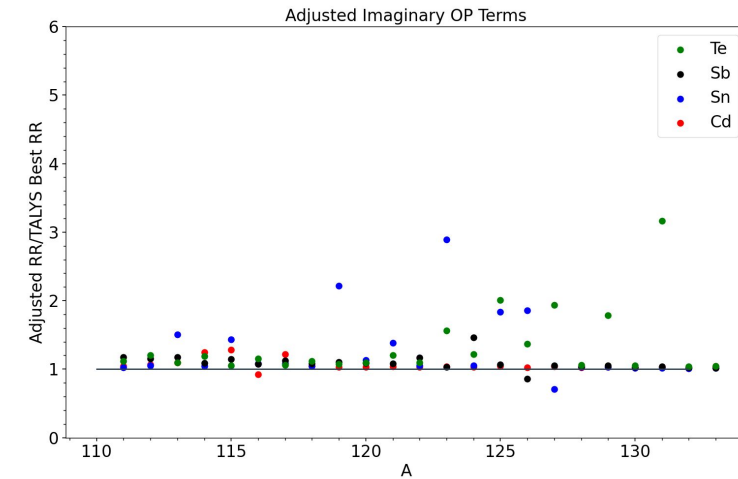
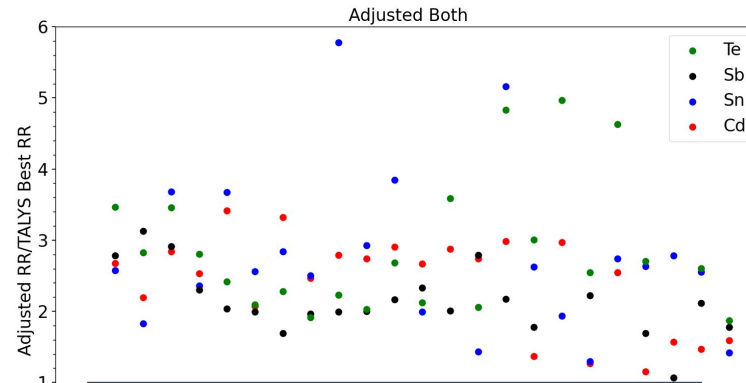
Astrophysical Reaction Rates (n, γ) at 30 keV

Rspincut=0.4



Preliminary

W1adjust n=2.5, W2adjust n=0.6



Improving reaction models can impact many different nuclear applications!

NSSC Sponsored Events:

Poster at:

- DOE NNSA University Program Review, *remote* (2021)

Other conferences, workshops, etc. attended w/ funding from NSSC:

Presented at:

- 8th Workshop on Level Density and Gamma Strength, *Oslo NO* (2022)
- 15th International Conference on Nuclear Data for Science and Technology, *remote* (2022)
- Low Energy Community Meeting, *East Lansing, MI* (2023)
- Taking the Temperature (T3) Workshop, *Athens, OH* (2023)
- 10th International Auger Symposium, *Montpellier FR* (2023)
- Radioisotope Production at SNS, *Oak Ridge, TN* (2023)

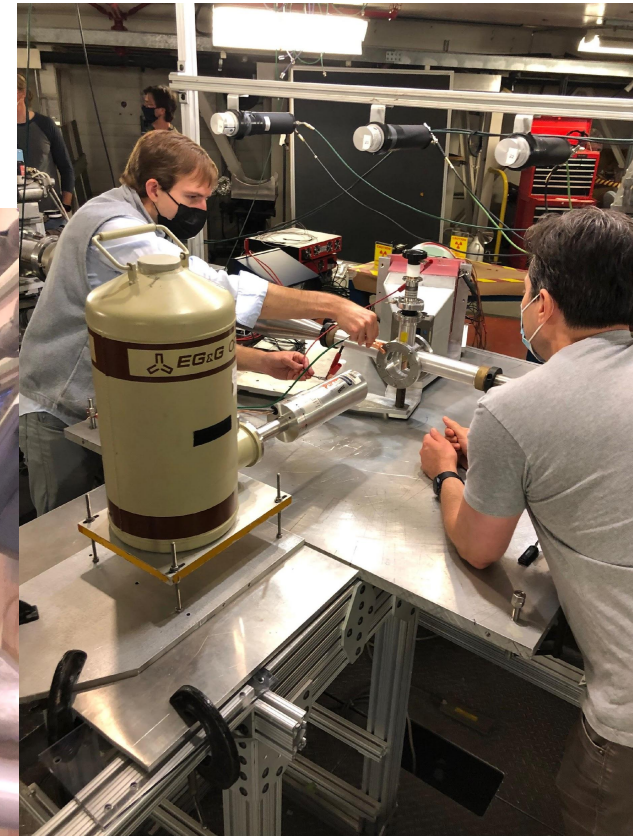
Poster at:

- Lawrence Livermore National Lab/Rutgers/UCB workshop, *Livermore CA* (2023)



Other research opportunities:

- Performed experiments at Los Alamos National Lab in 2020 and 2021
 - Toured Isotope Production Facility, Countroom, and Weapons Neutron Research facility
- Performed experiments at Brookhaven National Lab in 2021 and 2022
 - Toured Brookhaven LINAC Isotope Producer
- Performed experiments at Lawrence Berkeley National Lab in 2020 and 2022
 - Participated in other experimental work.
- Toured Facility for Rare Isotope Beams at MSU
- Wonderful opportunity to collaborate and network with folks from other universities and national laboratories



*Special thanks to the members of the
Tri-Lab Evaluated Data Collaboration:*

Lee Bernstein
Etienne Vermeulen
Dmitri Medvedev
Ellen O'Brien
Jon Batchelder
Eva Birnbaum
Cathy Cutler
Morgan Fox
Yun-Hsuan Lee
Jonathan Morrell
Meiring Nortier
Michael Skulski
Andrew Voyles

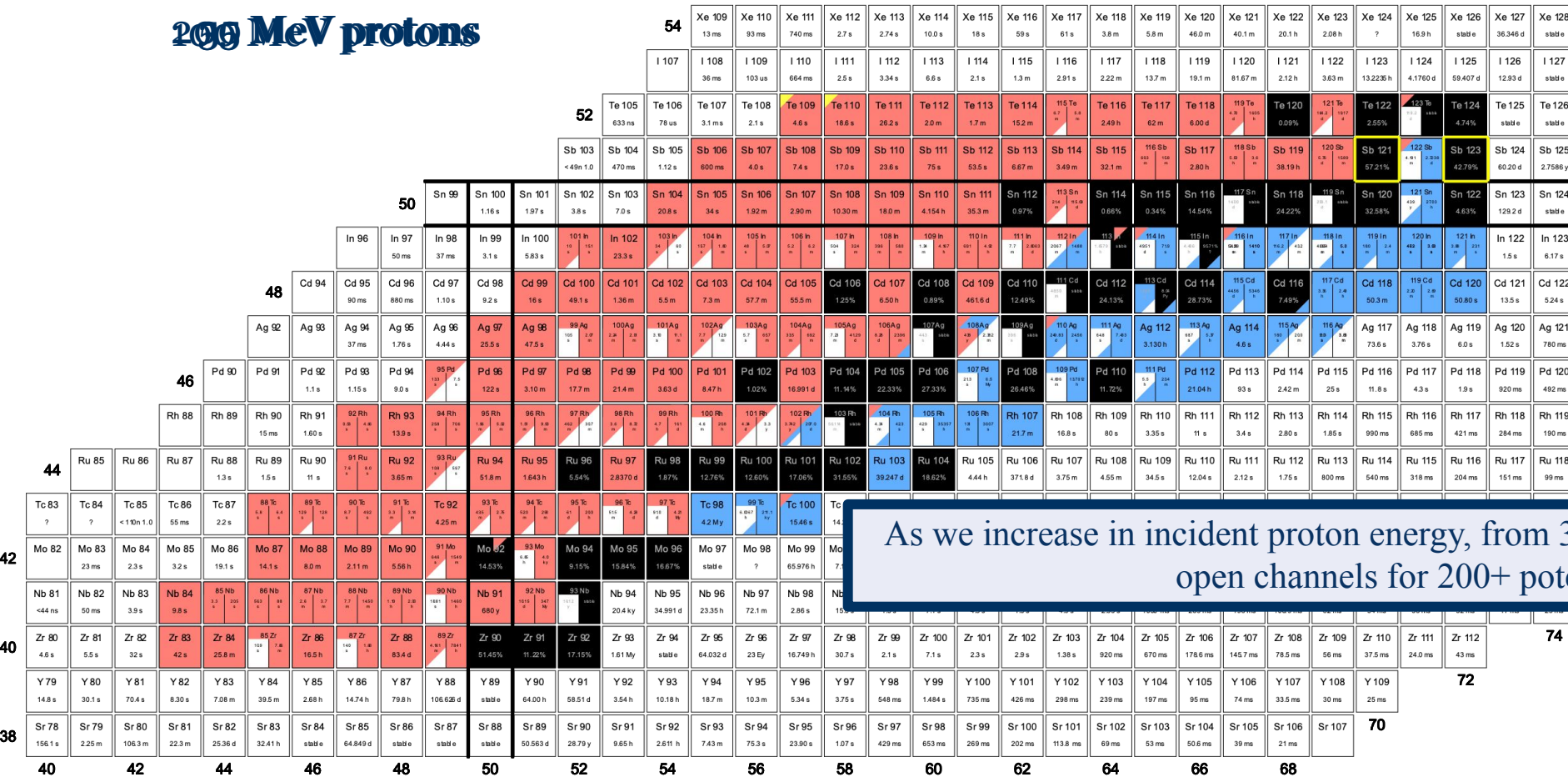


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Motivation

200 MeV protons



As we increase in incident proton energy, from 35 MeV all the way to 200 MeV, we open channels for 200+ potential products.